# Cosmic Exploration Caltech with Infrared Telescopes The Landscape Ahead

George Helou California Institute of Technology

Pasadena, May 2015

## Cosmic Exploration with Infrared Telescopes The Landscape Ahead

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- Facts: Missions & Themes
- Interpretation
- Questions

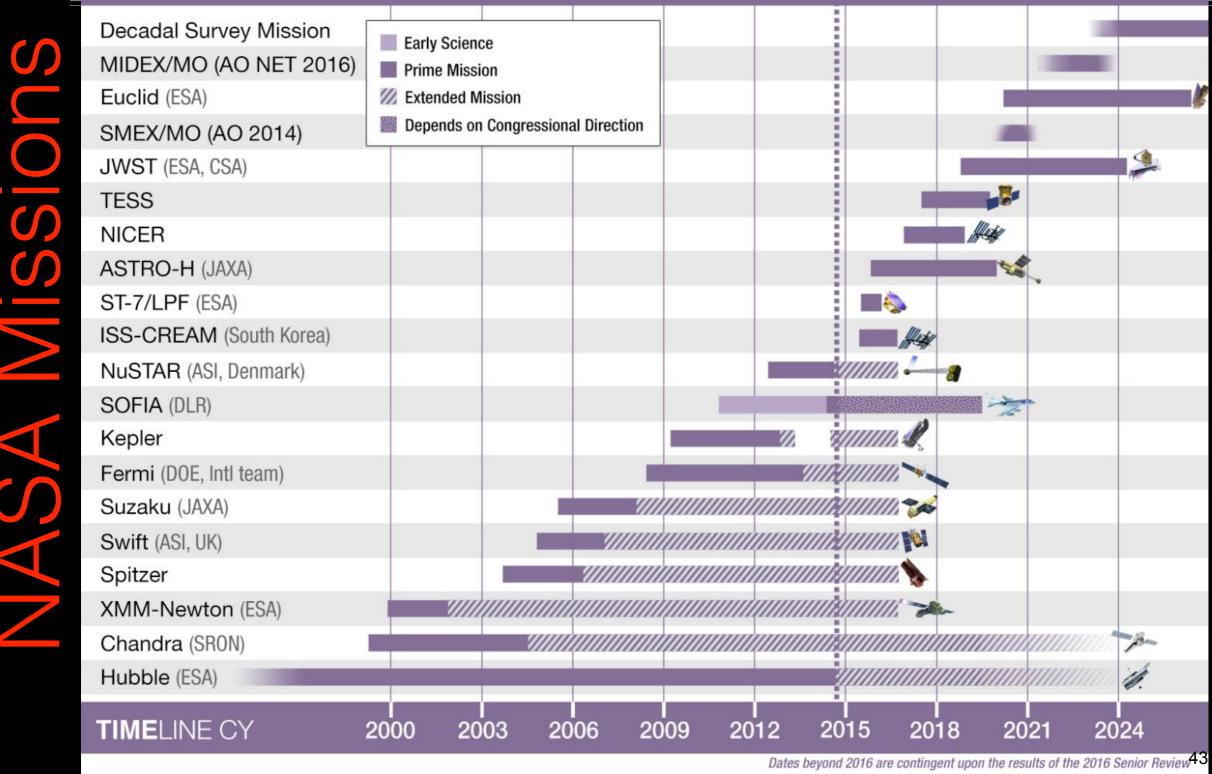
Pasadena, May 2015

#### • Facts: Missions & Themes

- Look ahead to 2020, 2025
- Interpretation
- Questions



#### **Astrophysics** Timeline



# Ground Ч Ч С Ц NASA,

#### Cosmic Dawn:

Reionization First stars, galaxies, BH

First stars, galaxies, BH Gastrophysics Physics of the Universe: CMB, LSS **Dark Matter & Energy** S C C C C C C C C C C C Gravity Waves, GR

# New Worlds:

Seeking habitable worlds Planetary System Formation

Solar System: Finding PHA (in time!) Our System's Dynamical History

hemes

S C C C C C C C C C C C C C

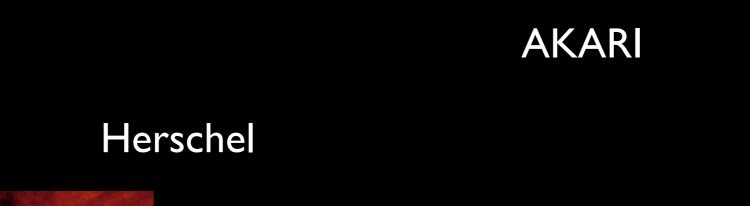
#### Where are we?



 Nel mezzo del cammin di nostra vita mi ritrovai per una selva oscura, ché la diritta via era smarrita.»
 — Dante Alighieri

Gustave Foré, La Selva Oscura

## The Infrared Space Telescope Legacy





ISO



IRAS



Spitzer

### The Next Steps

 Improved sensitivity, speed of mapping, or specialized instruments



• JWST

- WFIRST
- SPICA'

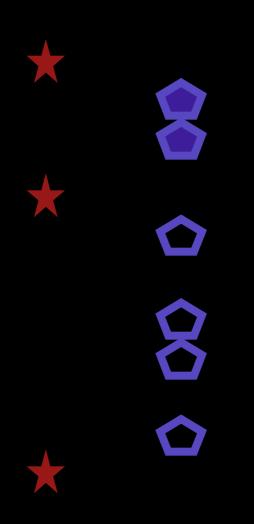
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#### Historical Perspective

- Past IR missions have been one of three types
  - Sky surveys (IRAS, AKARI, WISE, Planck)
  - Pointed observatories (ISO)
  - Pointed observatories with strong mapping capability (IRTS, Spitzer, Herschel, MSX)

 Future projects more clearly differentiated in spatial resolution and field of view

# Ground Ч Ч С S Ц NASA,



#### Wide-Field Surveys Capabilities

- Mostly in VIS (DES+, GAIA, LSST, Euclid), NIR (Euclid, WFIRST), radio (pre-SKA, SKA)
  - Euclid (VIS+NIR) ~24-25.5mag, z~[0.6-2.0]
  - ► WFIRST (NIR) ~26-27mag, z~[1.1-2.8]
    - Note: 25mag(AB) is 10<sup>-18</sup> W m<sup>-2</sup>

#### Wide-Field Surveys Science

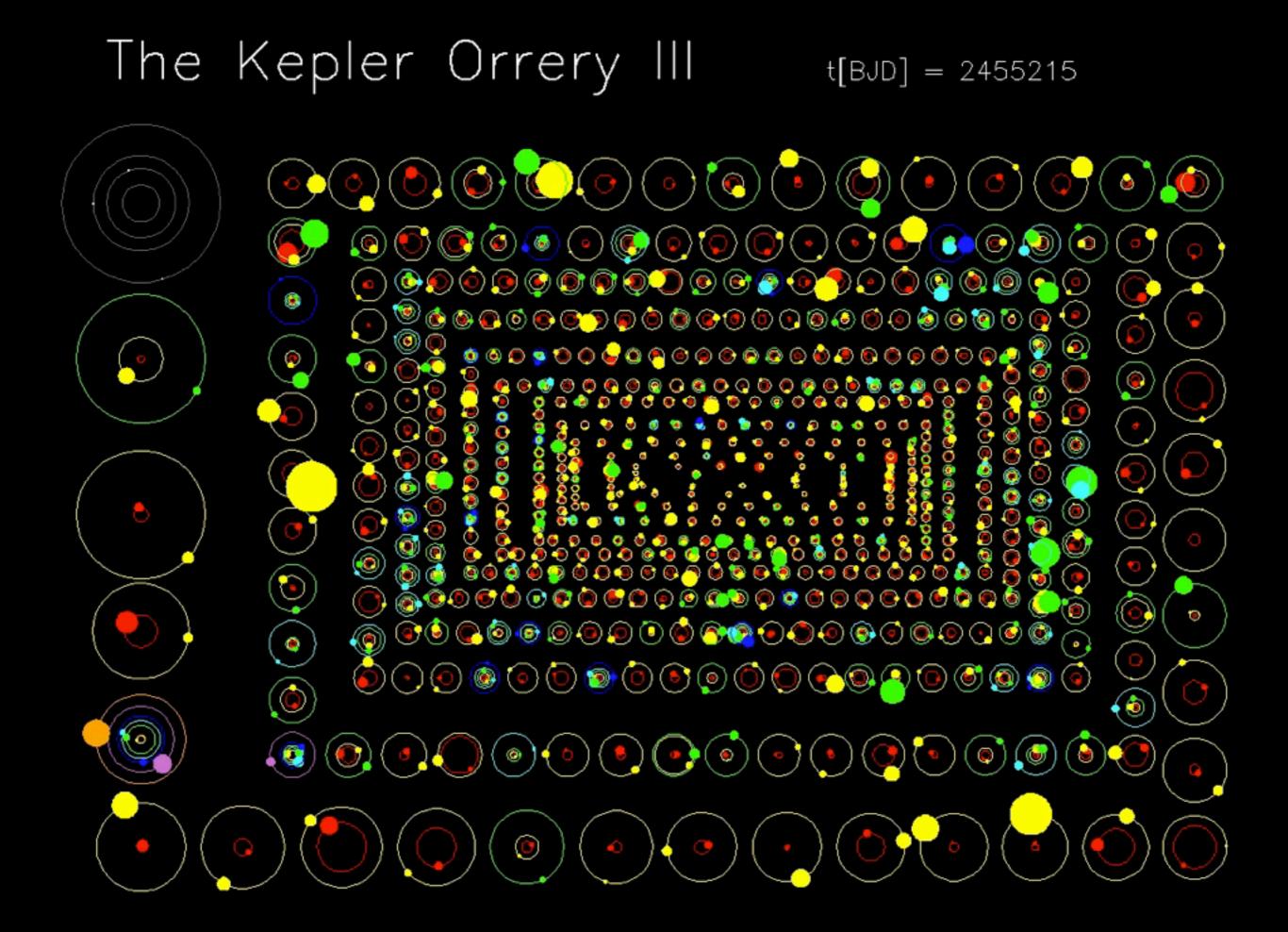
- Floods of interesting x-gal candidates calling for followup, thousands or more in a given category, with unknown MIR-FIR flux
  - Redshift range of peak SF, FIR background epoch
  - At FIR/VIS~I, these galaxies are CALISTO targets at f(100µm)≥100µJy and in some lines (>10<sup>-20</sup> Wm<sup>-2</sup>)
- Populations of stars in MW/Local Group, revealing outliers and missing links,GAIA: amazing improvement of MW/LG measure

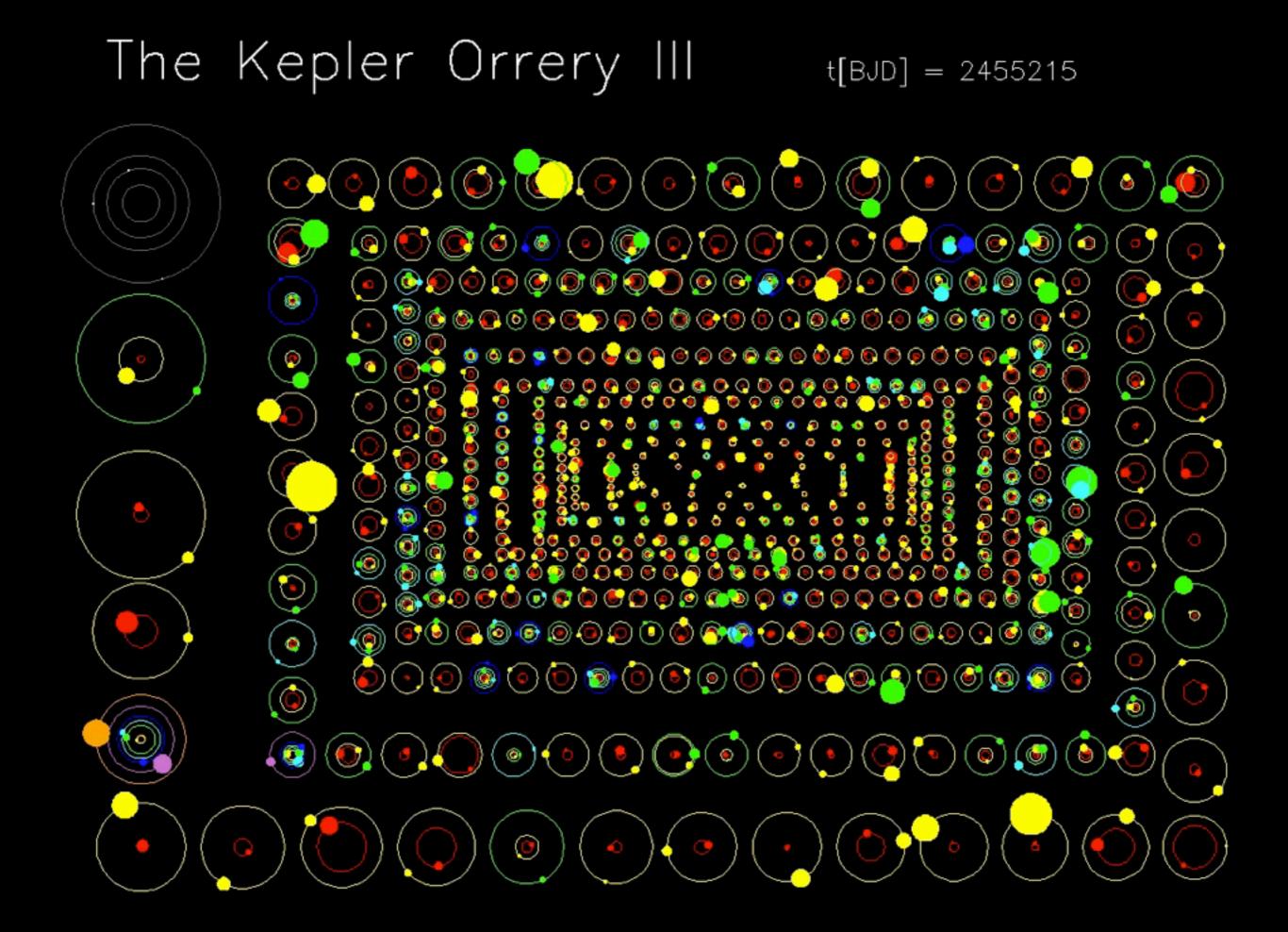
#### **High-Resolution Telescopes**

- ALMA will be unbeatable at >300µm
  - Revealing intricate details (down to 0.007") of sources at all distances, Solar System to z>8
- JWST will dominate at <30µm, including deep pencil-beam surveys</li>
  - Probing the x-gal universe at z>4 (0.07" at  $2\mu$ m)
- GSMT in 20s, similar capabilities ~0.005"
- SPIRIT offers ~0.2" at 60µm (close enough)

- Do we explore the same landscape as the wide-field surveys, followup those objects?
  - Most unlikely to find a redundant universe VIS/FIR
  - The 20s survey suite is incomplete without FIR
- Do we go for high spatial resolution?
  Complement ALMA, JWST, GSMT with a FIR interferometer?

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- Do we go for high spatial resolution?
  Complement ALMA, JWST, GSMT with a FIR interferometer?
- In any case, let's not ignore topics "not well suited to FIR", and let's think big picture





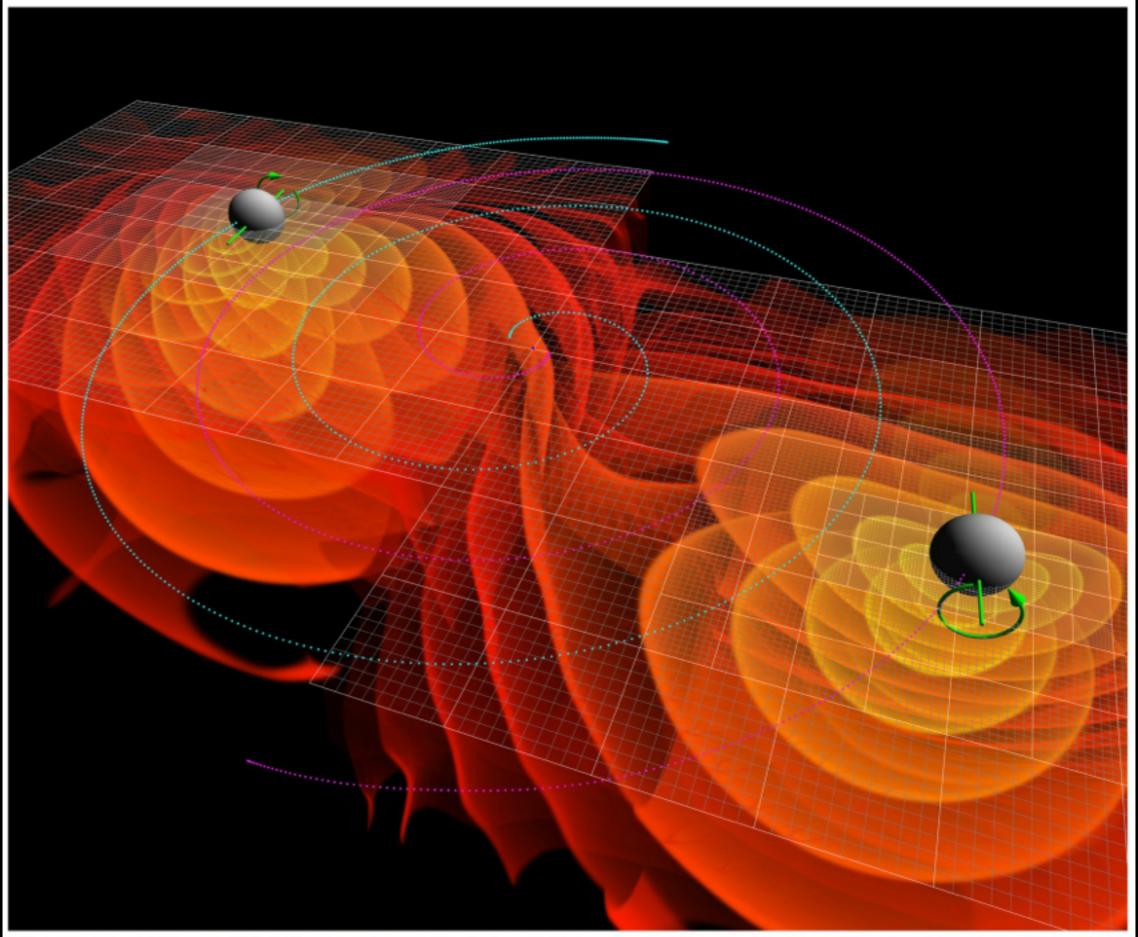


Figure 4.7 This snapshot from a numerical simulation shows the gravitational waves produced by a pair of merging black holes. Credit: Chris Henze (NASA Ames), from a simulation by J. Centrella, B. Kelly, J. Van Meter, and J. Baker (NASA GSFC)



Figure 1.2 Chart of the missions currently planned for launch during the Near-Term Era and of the notional missions of this roadmap for the Formative and Visionary Eras.

#### The Programmatic Landscape

#### The Infrared Landscape

#### The Infrared Landscape



